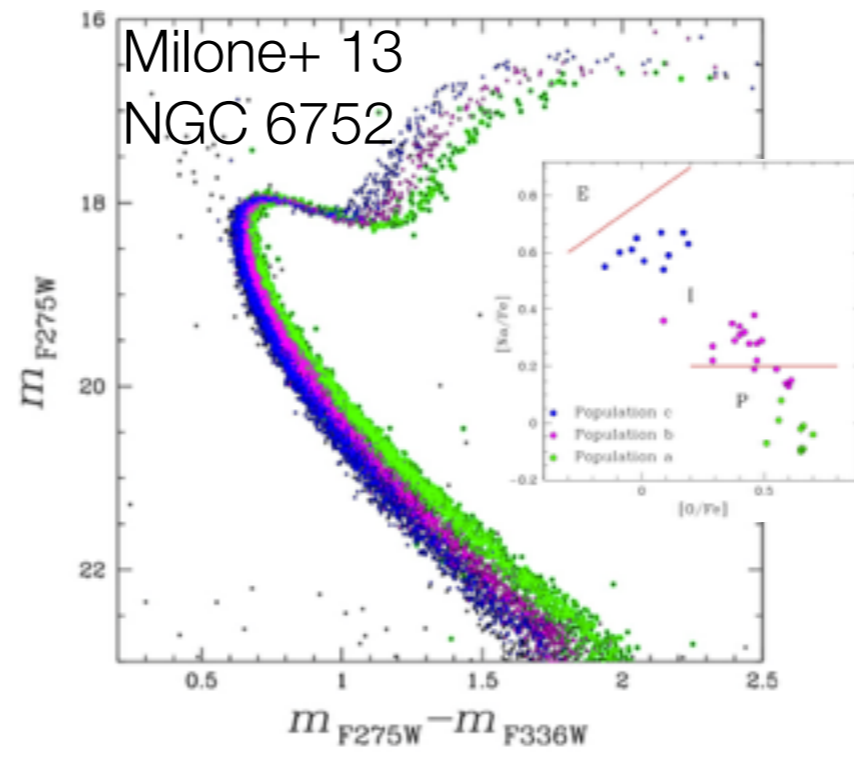


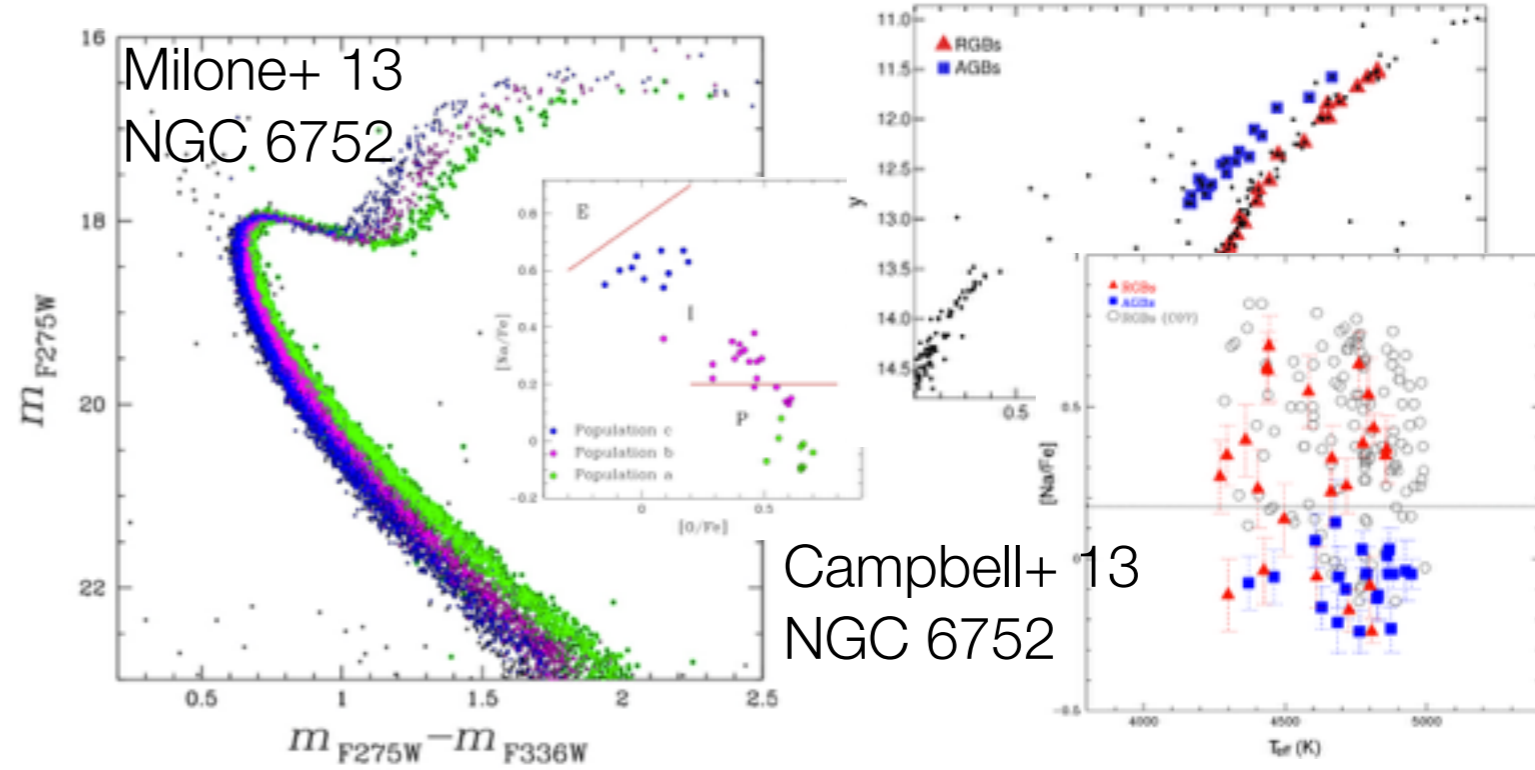
Chemical properties of AGB stars in GCs

I. Cabrera-Ziri, W. Chantereau, I. Claydon, R. Cohen, B. Dias, L. Grossler, I. Jean-Baptiste, G. Nandakumar, M. Powalka

Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).



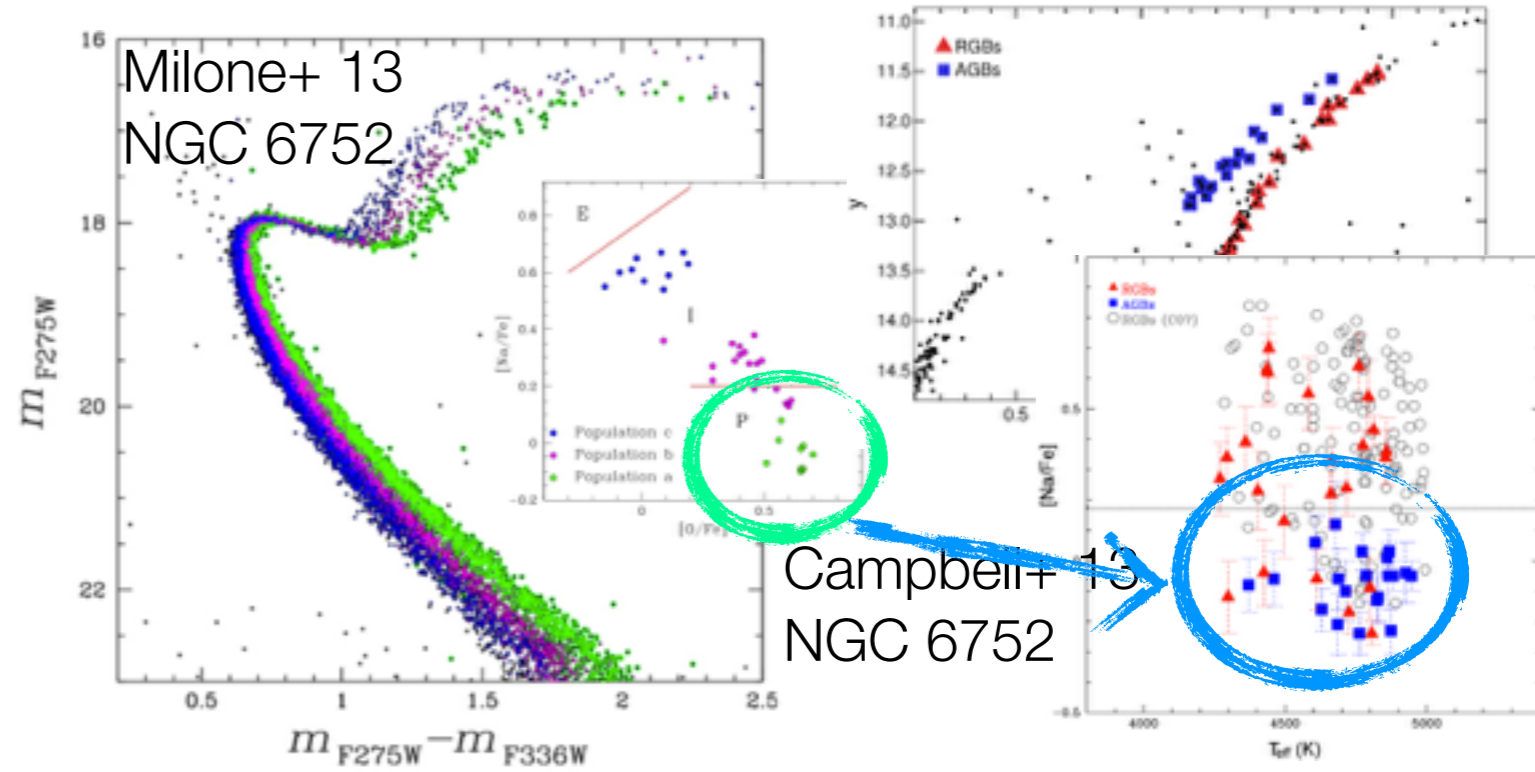
Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).



AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.

Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).

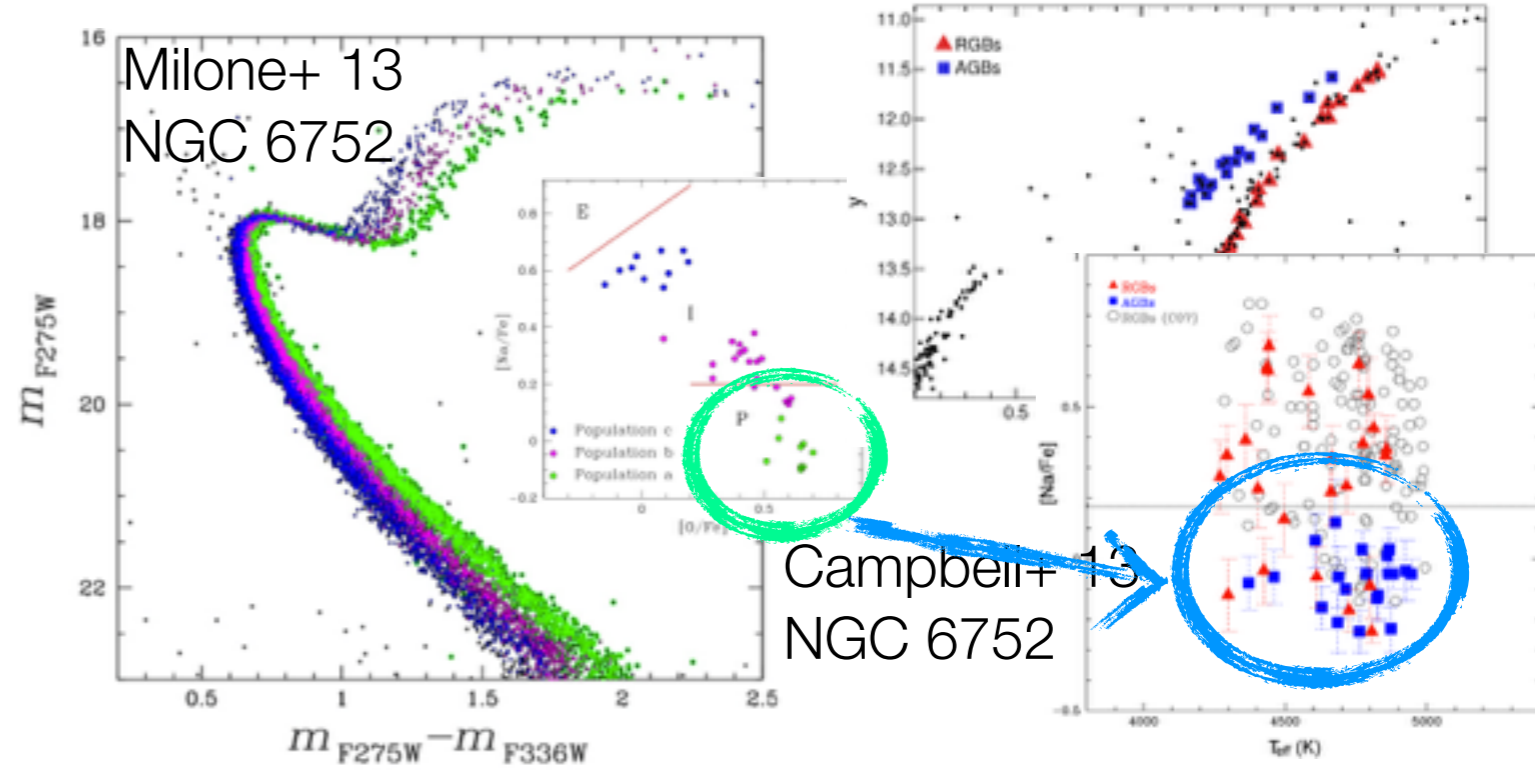
The AGBs of some GCs seems to be missing the enriched pop., **i.e. 70% of the GC stars!!!**.



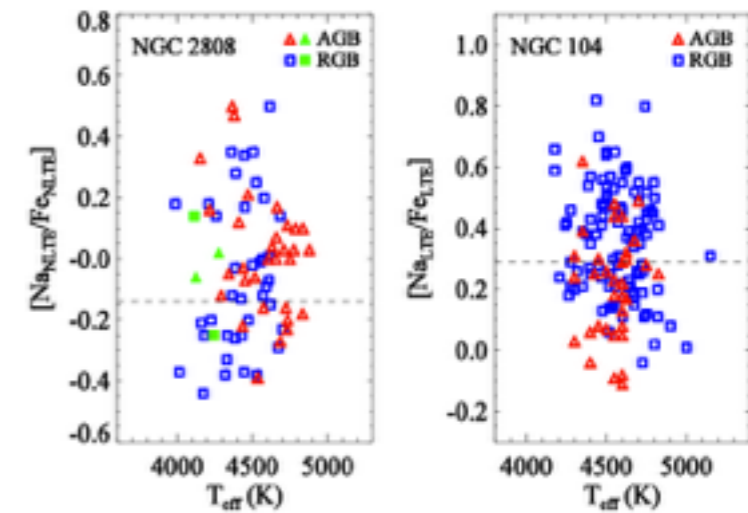
AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.

Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).

The AGBs of some GCs seems to be missing the enriched pop., **i.e. 70% of the GC stars!!!**



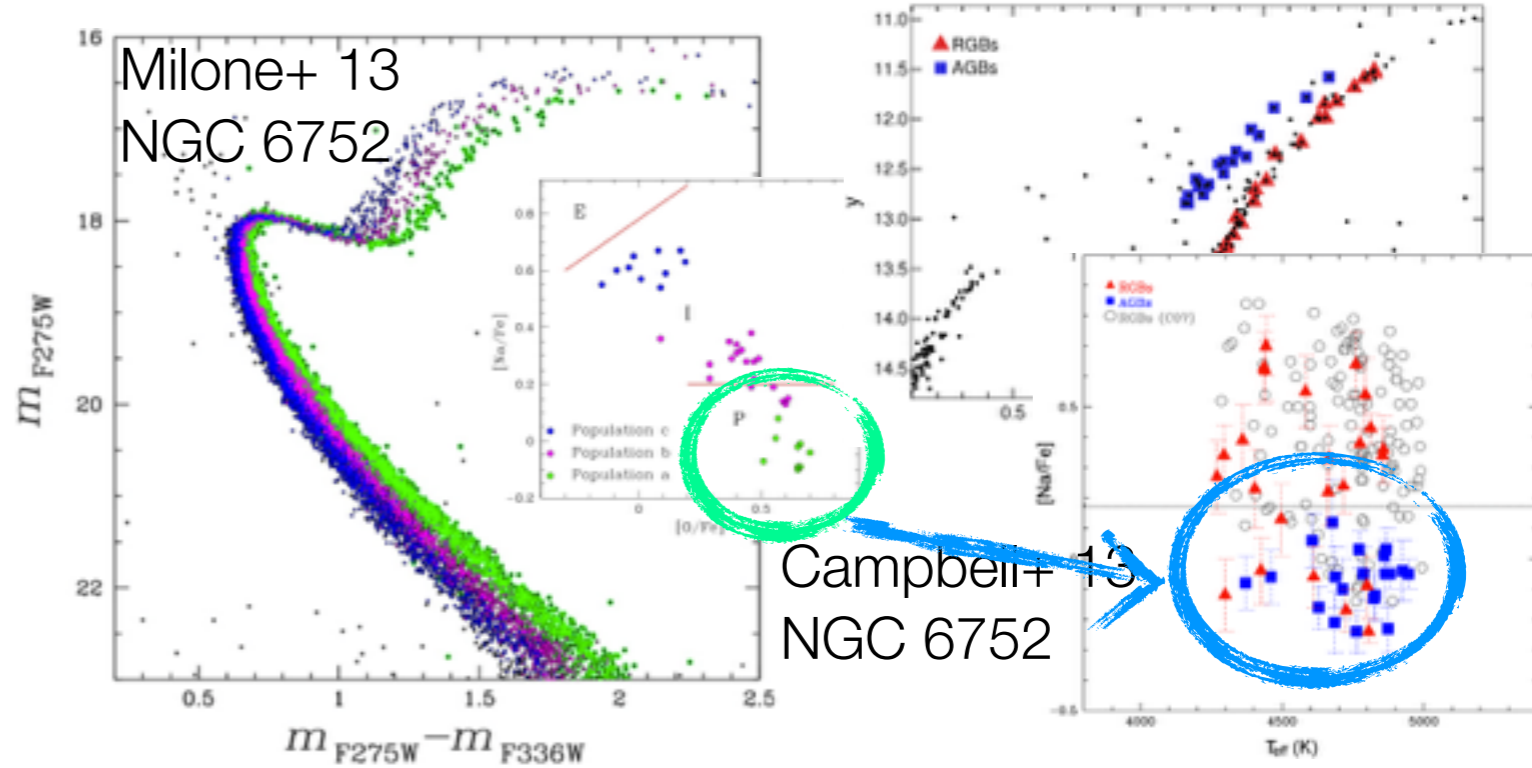
AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.



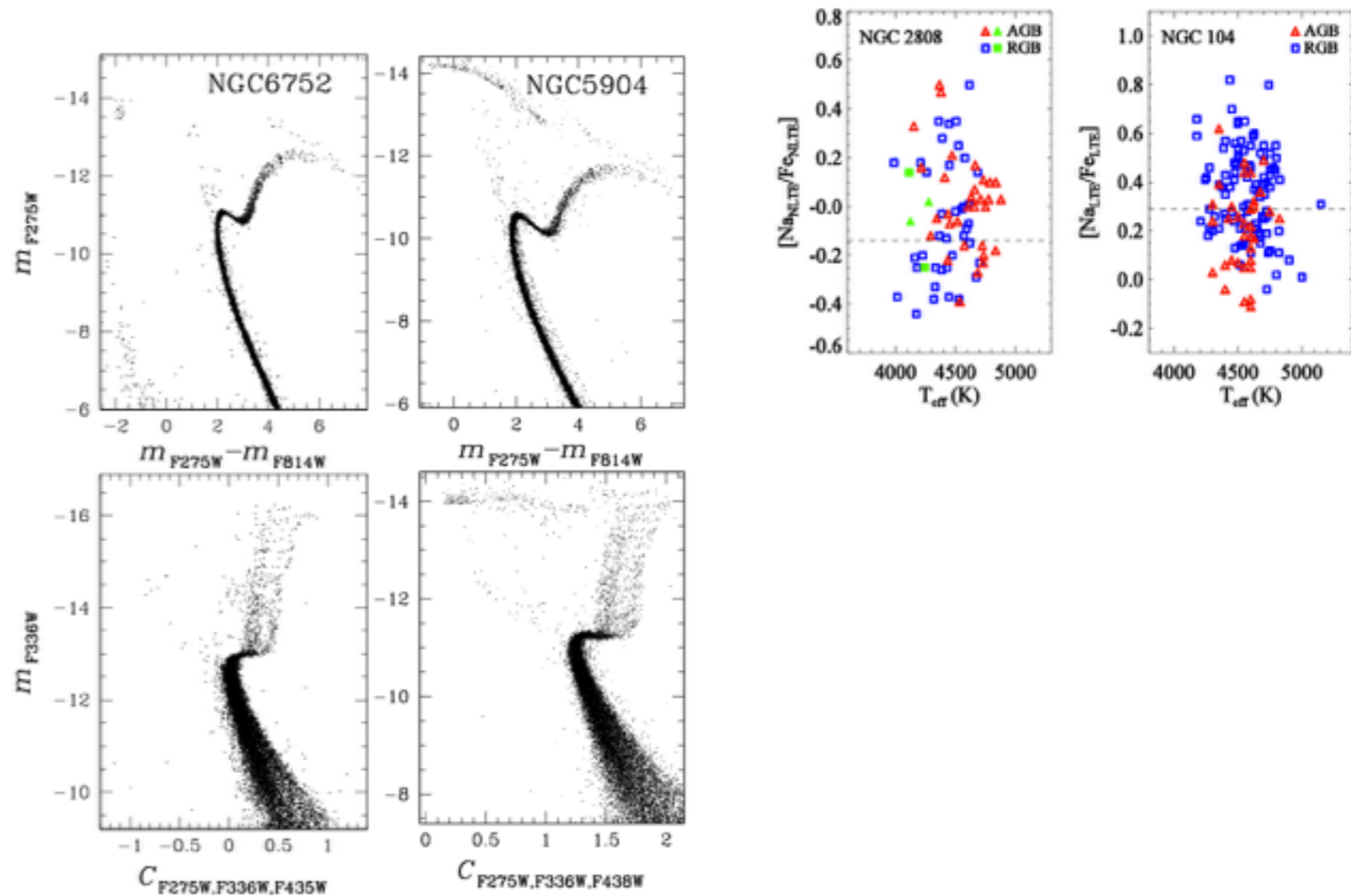
Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).

The AGBs of some GCs seems to be missing the enriched pop., **i.e. 70% of the GC stars!!!**

HST Treasury Large Programme (60 GCs) observed with magic filters.



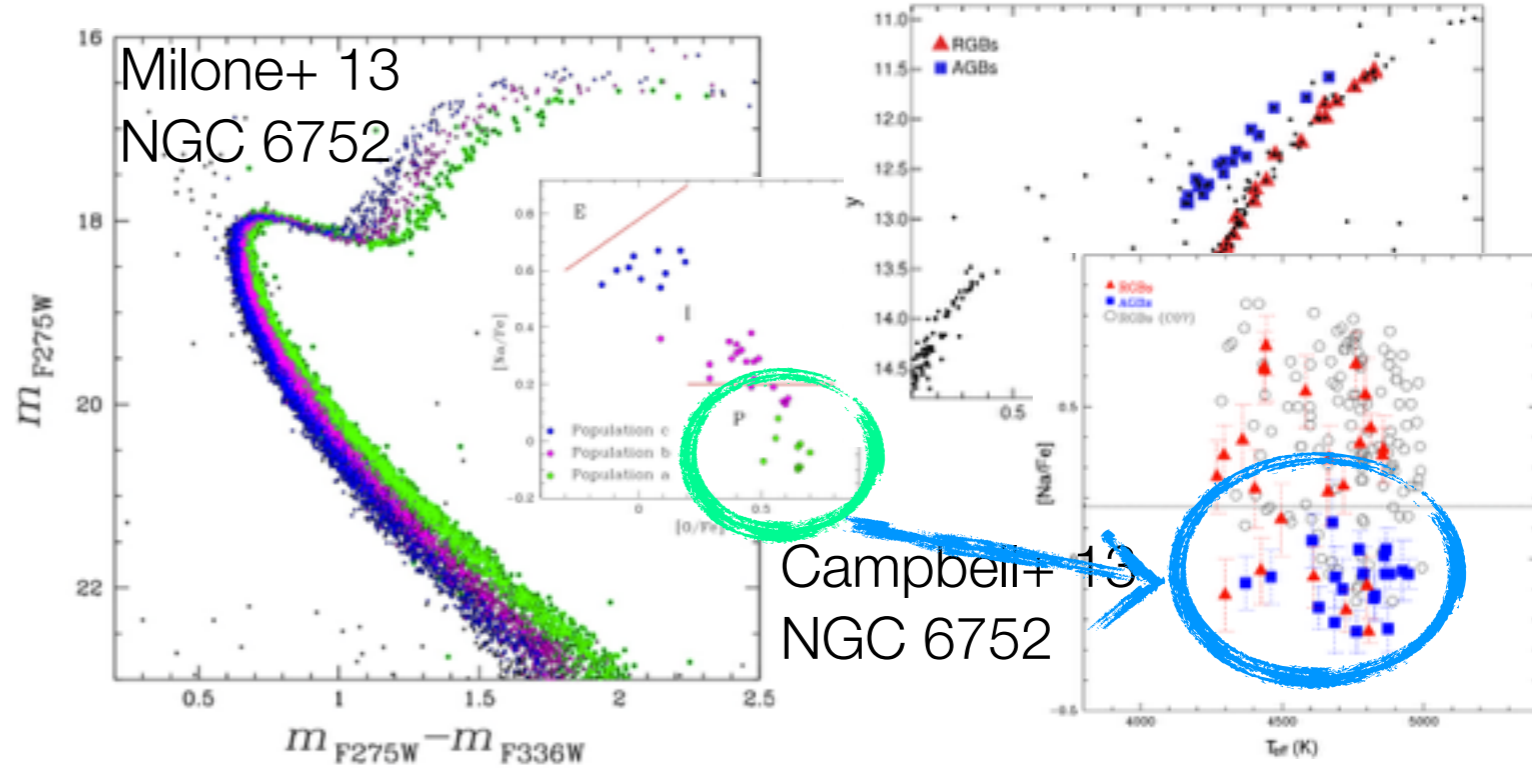
AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.



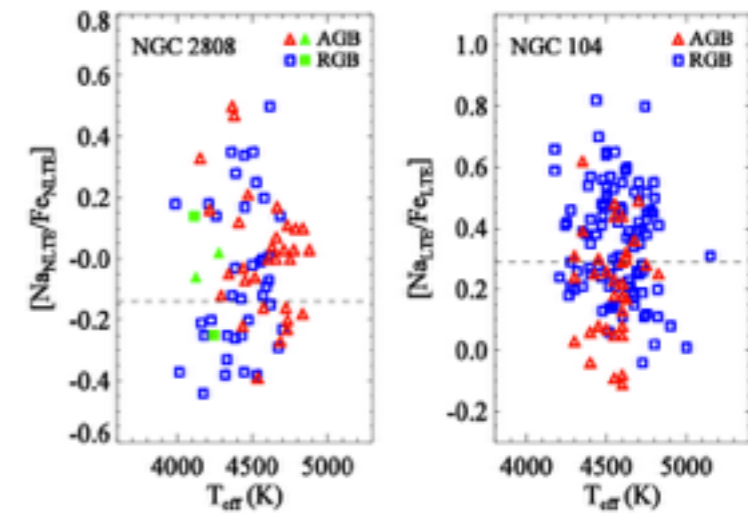
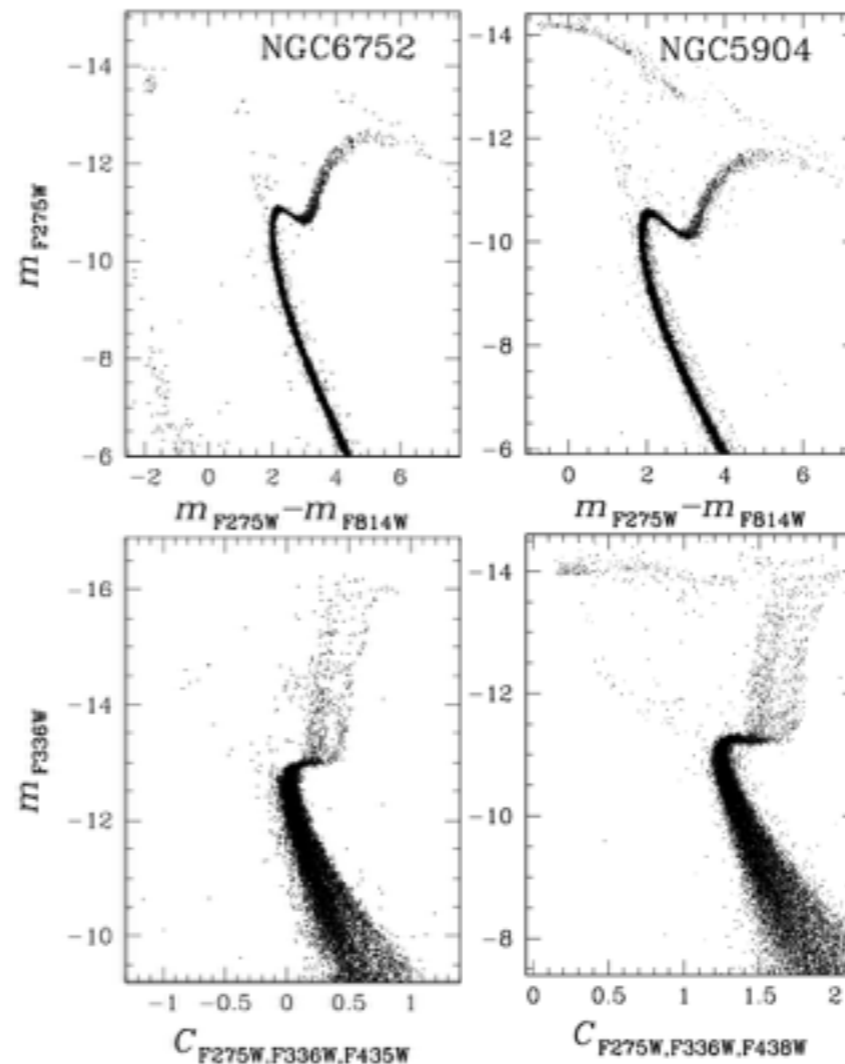
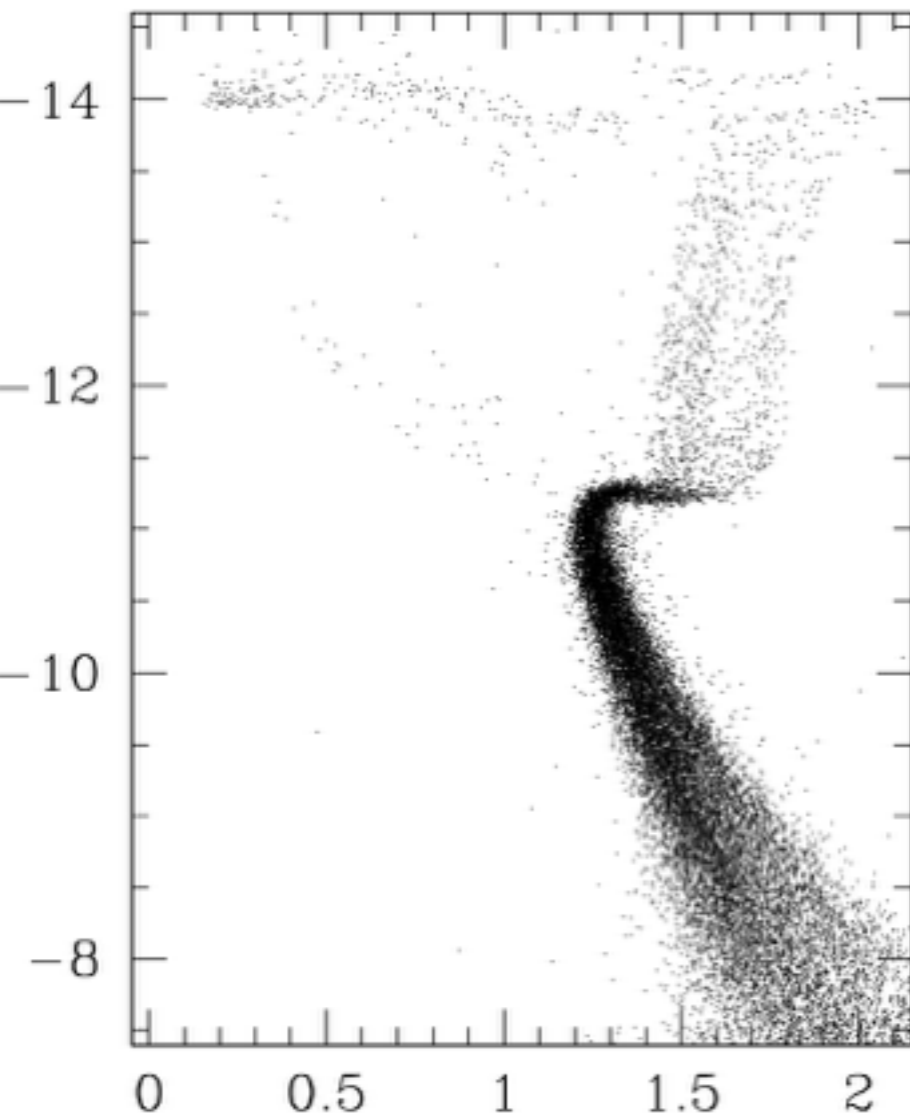
Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).

The AGBs of some GCs seems to be missing the enriched pop., **i.e. 70% of the GC stars!!!**

HST Treasury Large Programme (60 GCs) observed with magic filters.



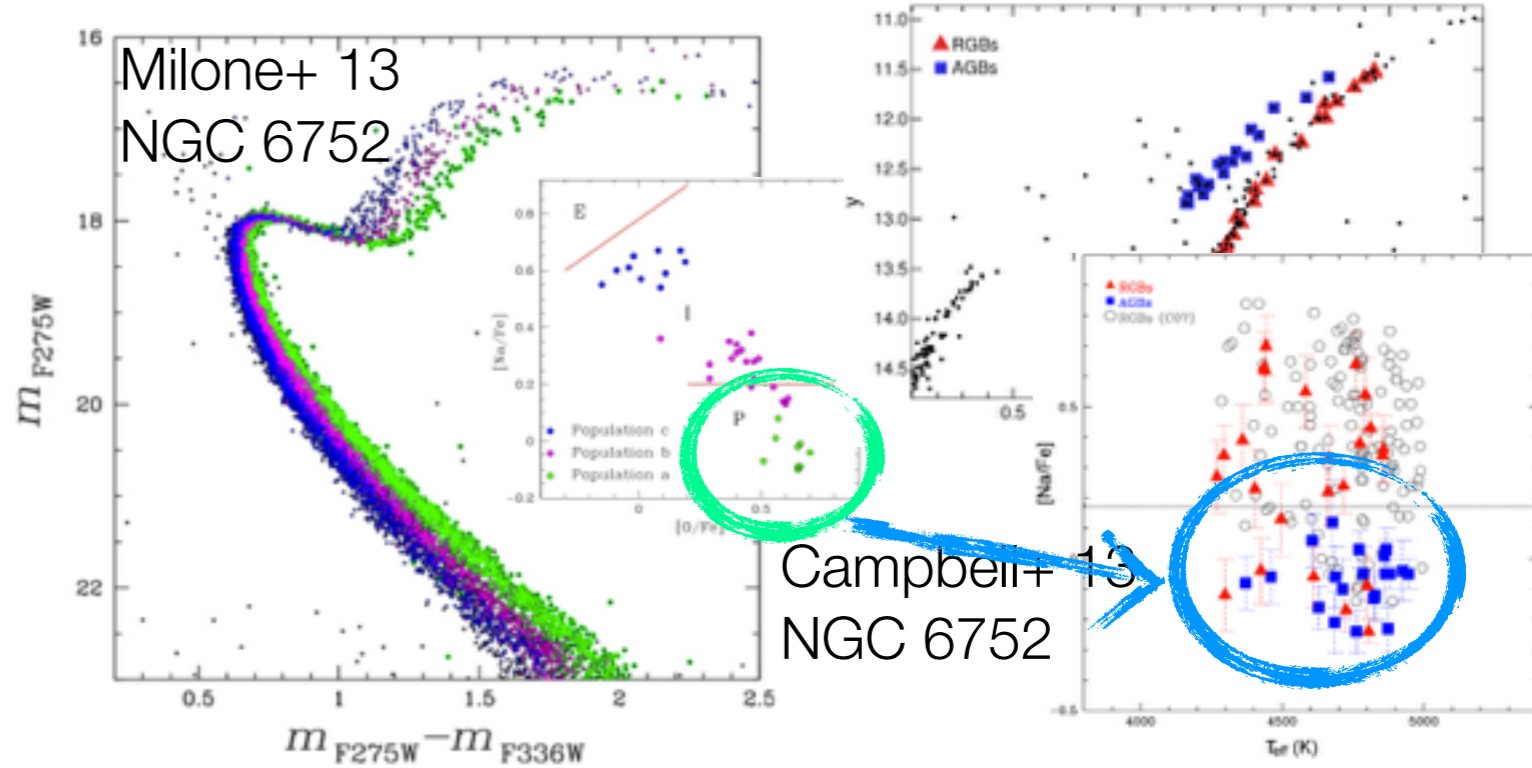
AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.



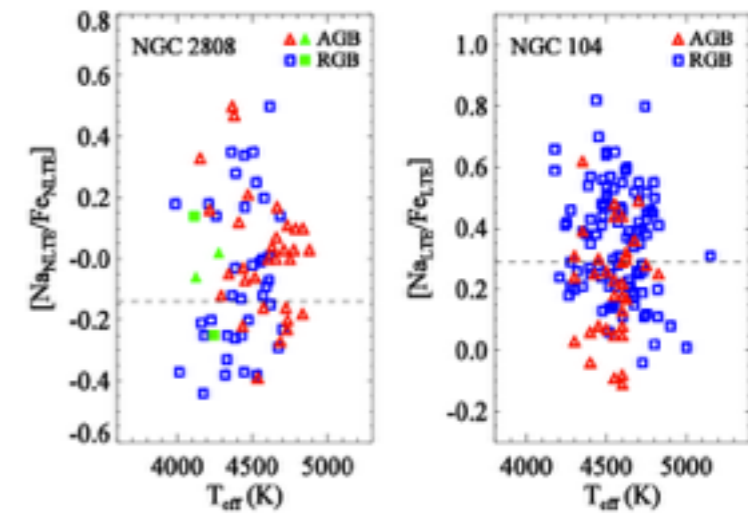
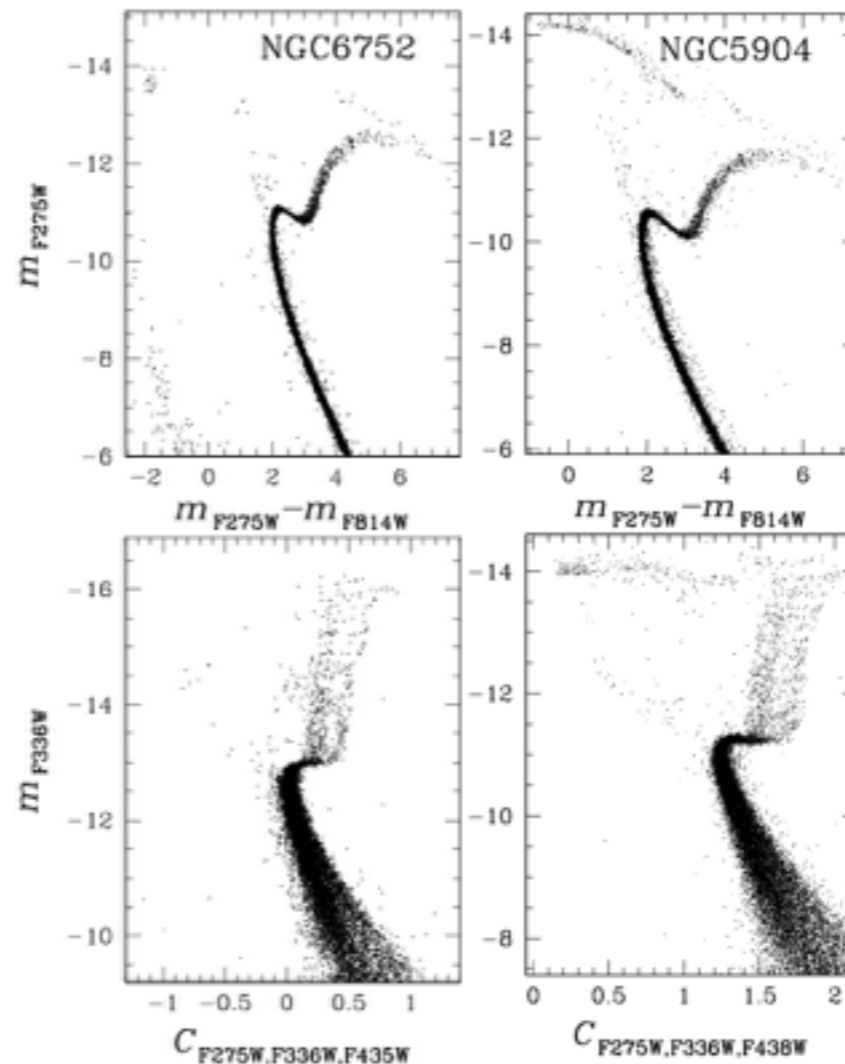
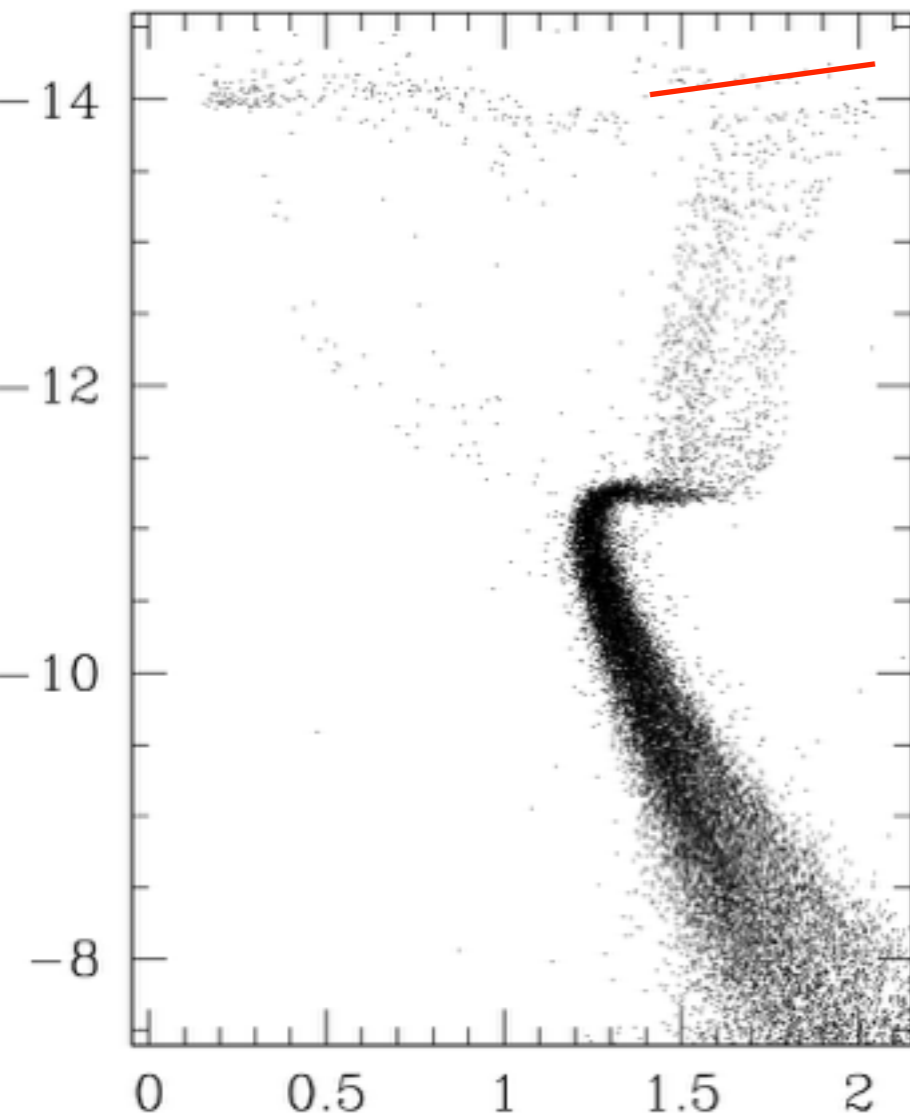
Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).

The AGBs of some GCs seems to be missing the enriched pop., **i.e. 70% of the GC stars!!!**

HST Treasury Large Programme (60 GCs) observed with magic filters.



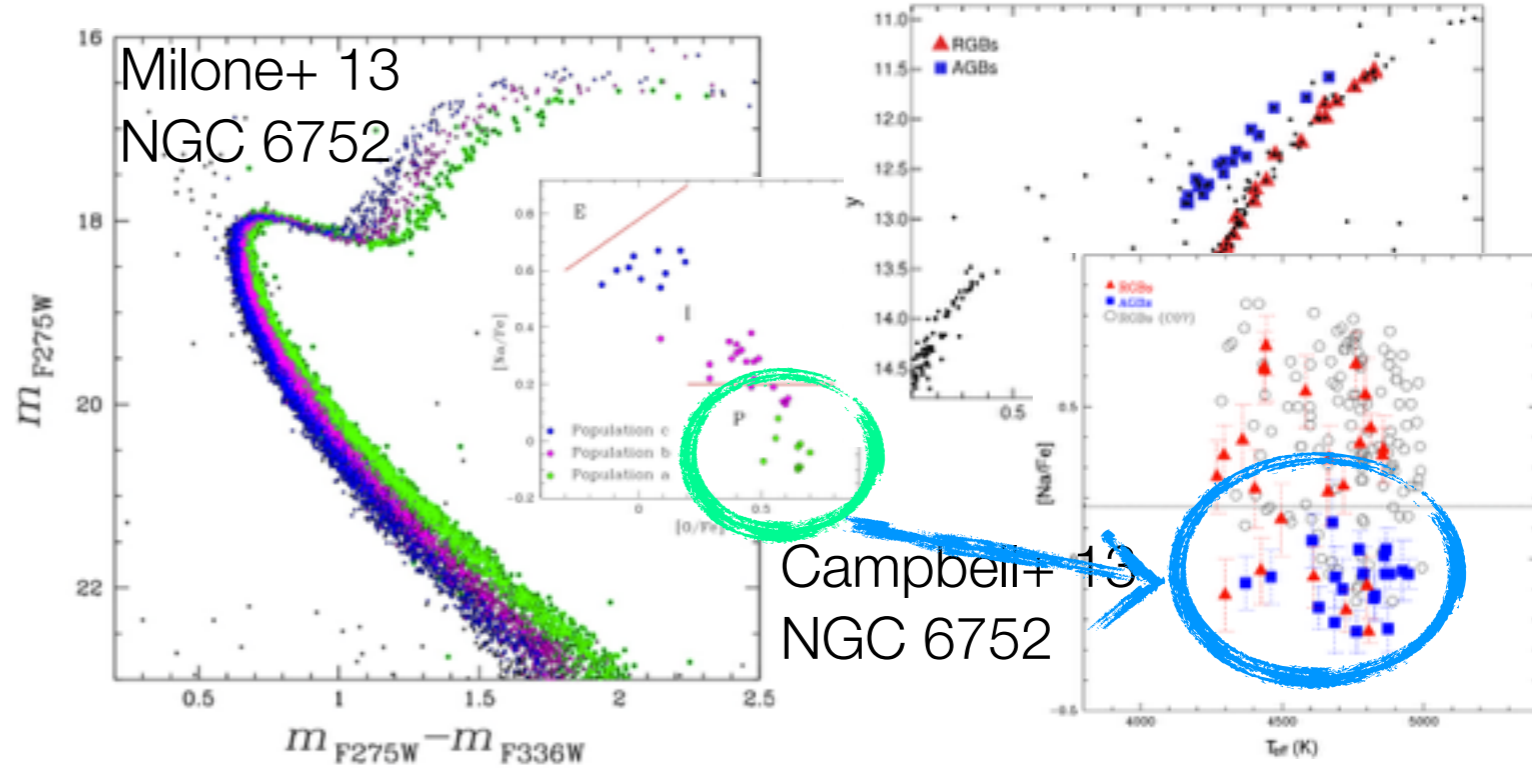
AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.



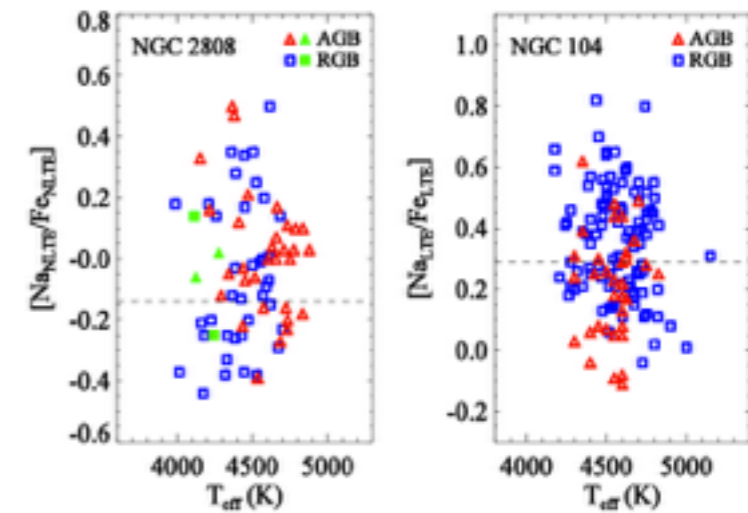
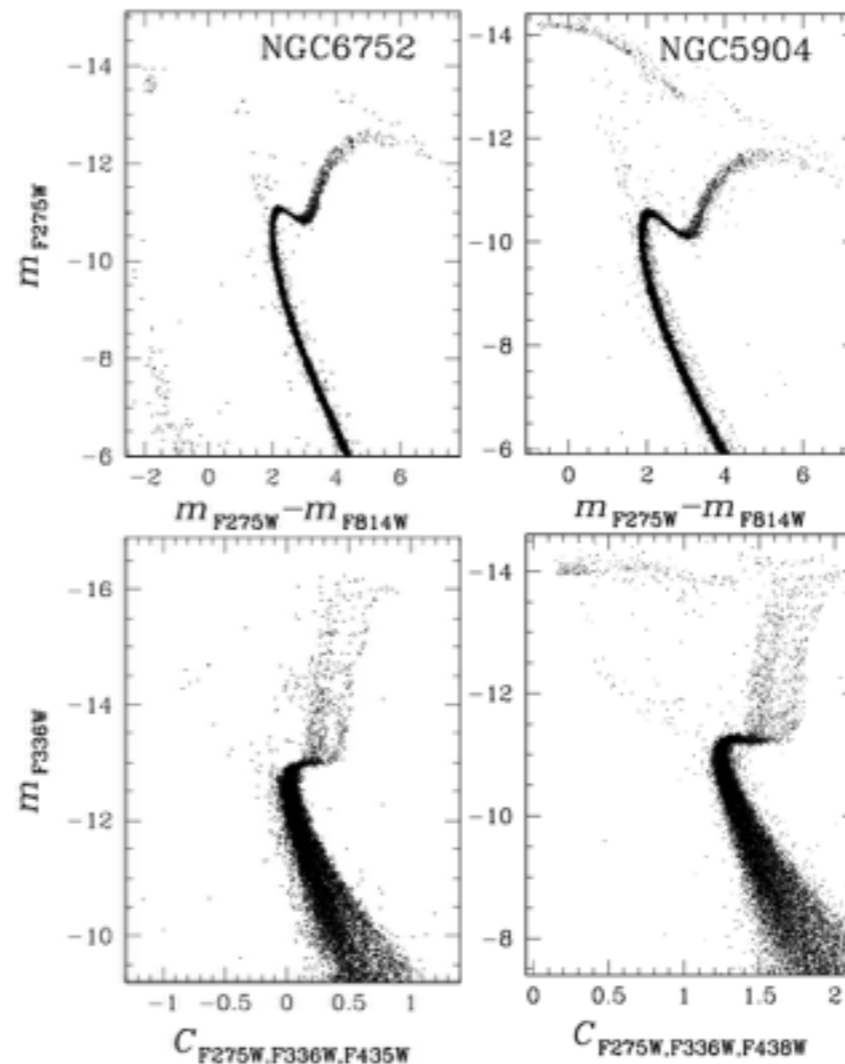
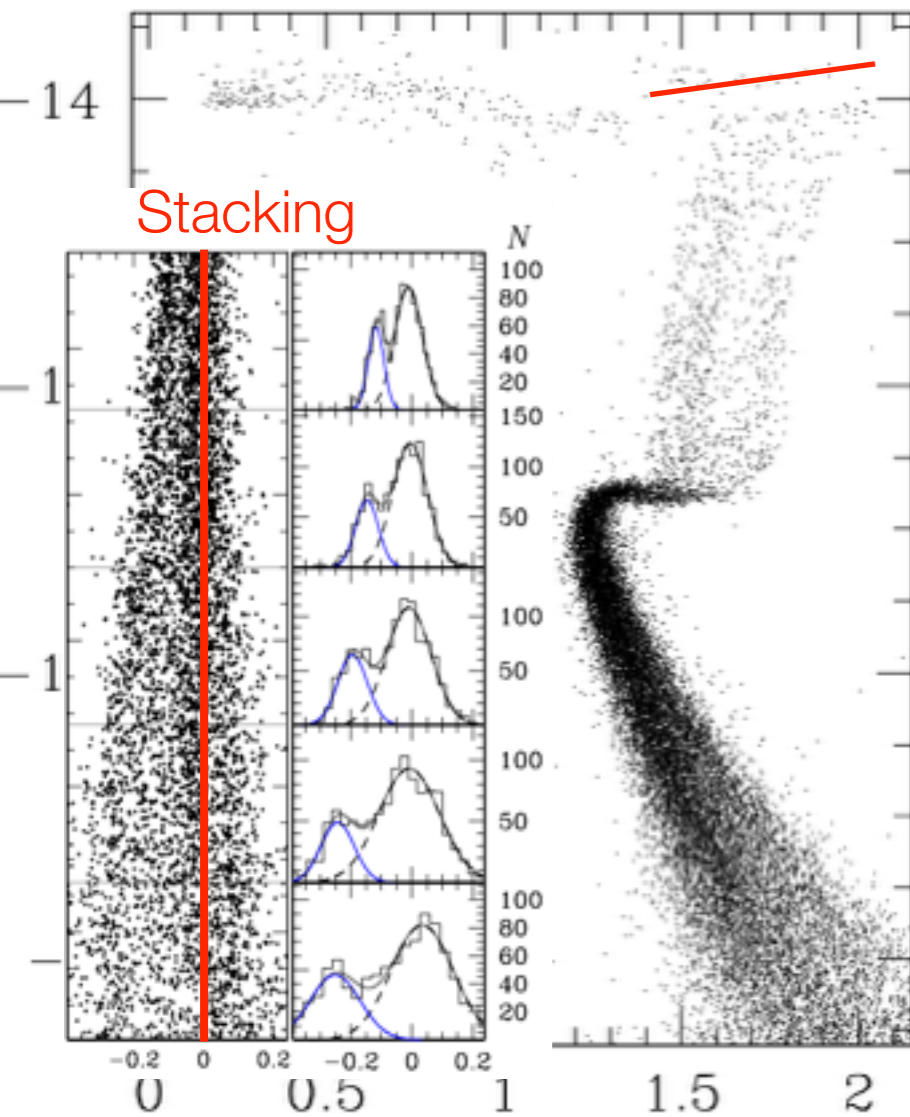
Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).

The AGBs of some GCs seems to be missing the enriched pop., **i.e. 70% of the GC stars!!!**

HST Treasury Large Programme (60 GCs) observed with magic filters.



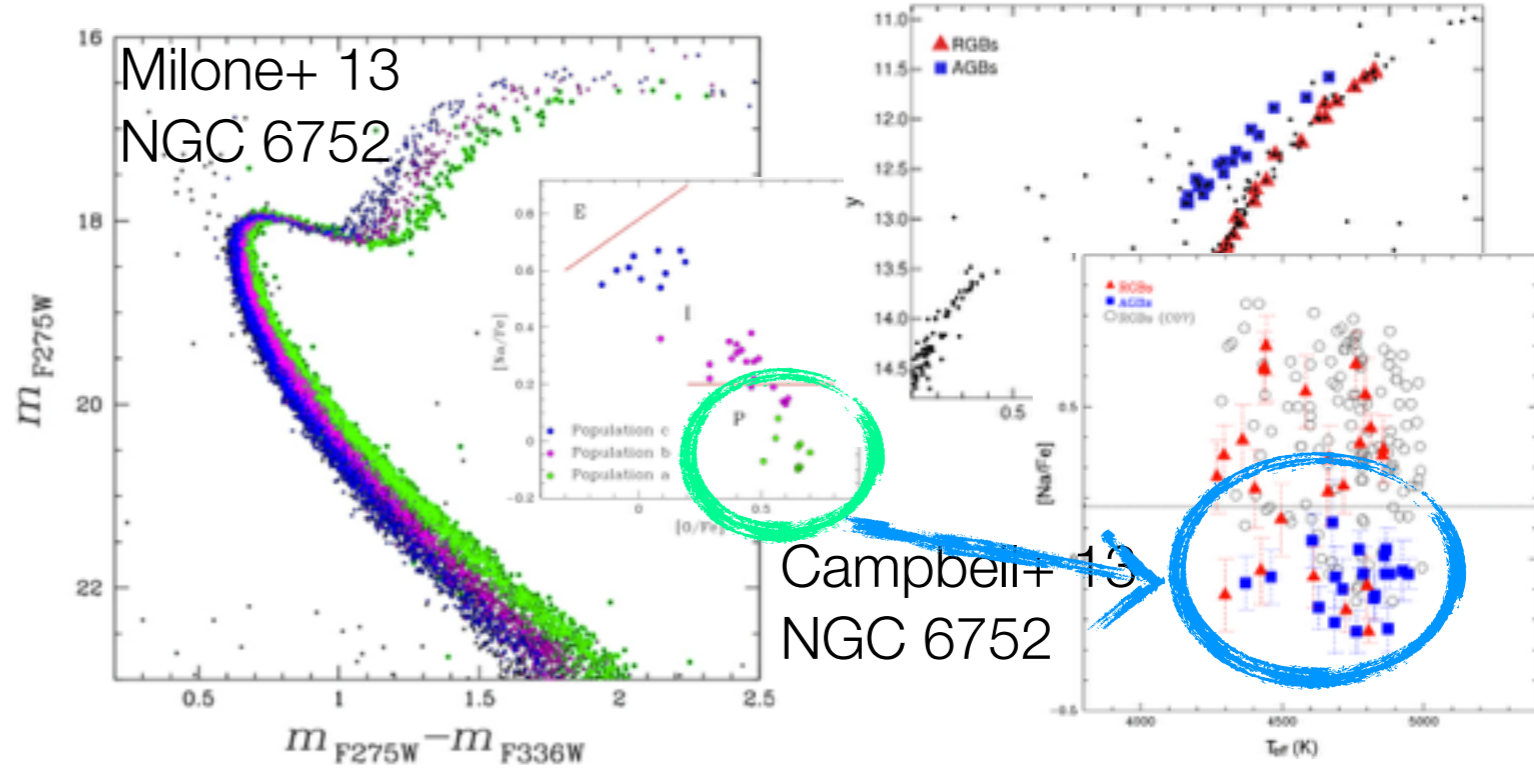
AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.



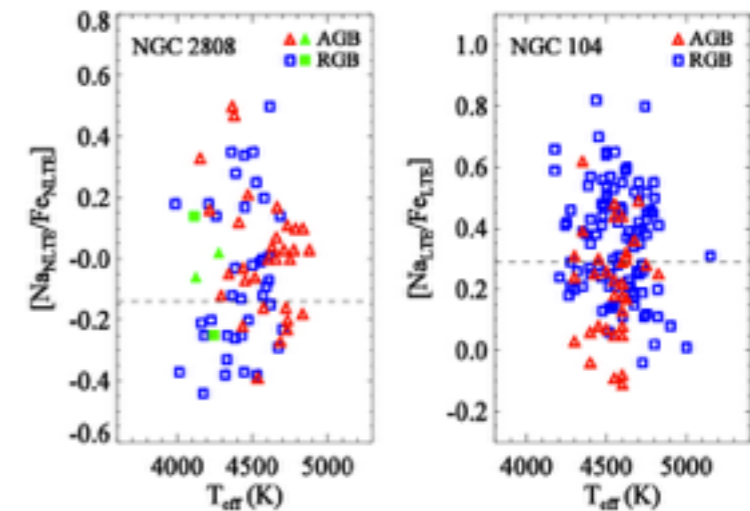
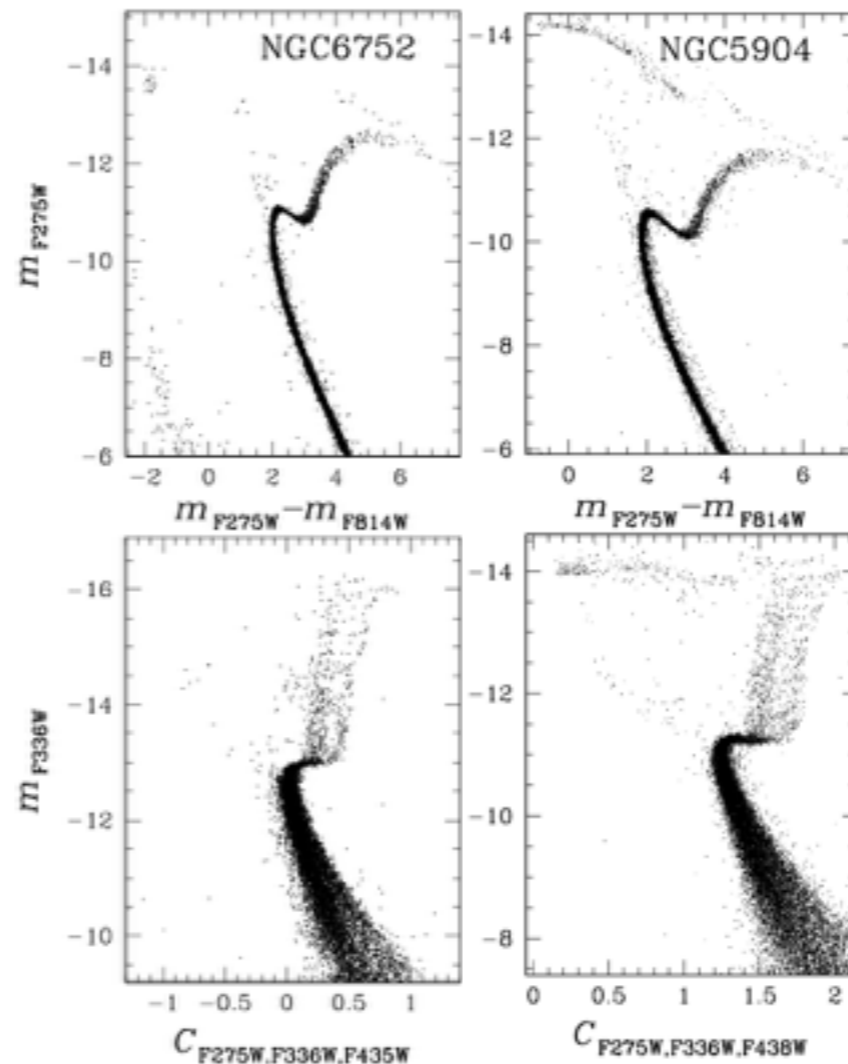
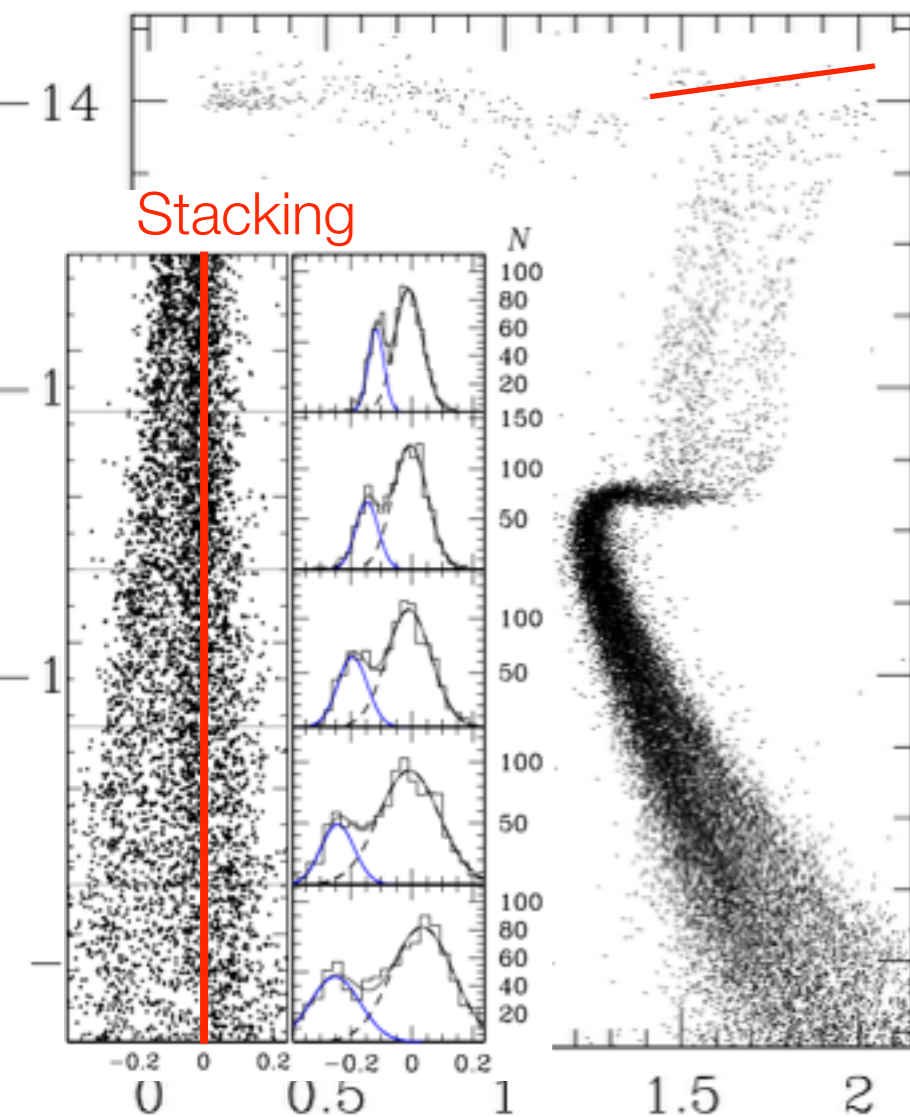
Multiple stellar populations (MP) have been found in all evolutionary phases of GCs (e.g. from MS to HB).

The AGBs of some GCs seems to be missing the enriched pop., **i.e. 70% of the GC stars!!!**

HST Treasury Large Programme (60 GCs) observed with magic filters.



AGB: **NO** CN-strong pop. (Norris+ 81) & **NO** Na rich pop.



Already at this early stage we would have a lot of information of CN-strong stars, i.e. possible polluters, fraction, positions, enrichment range, etc, etc...

Immediate objectives

- 47/60 GCs accessible from VLT: Propose a large programme on UVES-FLAMES @ VLT to get HR spectra ($R \sim 40k$) of complete population of AGB stars in GCs.
- Create evolutionary tracks for He-enriched AGB stars.
- Build synthetic libraries for abundance measurements.
- Run N-body simulation with stellar evolution (of enriched and pristine chemistry) in GCs to match observed properties of the Treasury sample and AGB properties.

Working groups + timescales = publications

- Observing:

- Candidate selection (few months, HST archive data)
- Proposal (1yr), observations (1yr), data reduction/analysis (1yr)

Roger Cohen
Ingrid Jean-Baptiste
Lucas Grosset
Bruno Dias
Ivan Cabrera-Ziri

- Theory:

- Multiple pops. of GCs scenarios (1yr)
- Stellar Evolutionary tracks for AGB (1yr)

William Chantereau
Ivan Cabrera-Ziri
Mathieu Powalka

- Modelling:

- Synthetic spectral libraries of AGB stars: NLTE/3D (1yr)
- N-body/Dynamics of multiple pops of GCs (1yr)
- SSP models with empirical library from the VLT survey (months)

Bruno Dias
Mathieu Powalka
Ian Claydon
Lucas Grosset
Govind Nandakumar

Publications (at least 7 to start)

- Chemistry from CMD (stack of AGBs from HST CMDs)
- Chemistry from abundances using NLTE/3D models (Na, O, Mg, Al at least for AGBs)
- Evolutionary tracks AGB phase: 2 partial papers while it runs within 1 yr
- Empirical spectral library of AGBs (reduced, calibrated spectra = public)
- Dynamics vs. observations (does AGB has a special dynamical evolution?)
- SSP models updated with known AGB spectra.